

Double nuclear structure discovered in ${}^3\text{C}84$

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EVN 2018, Granada

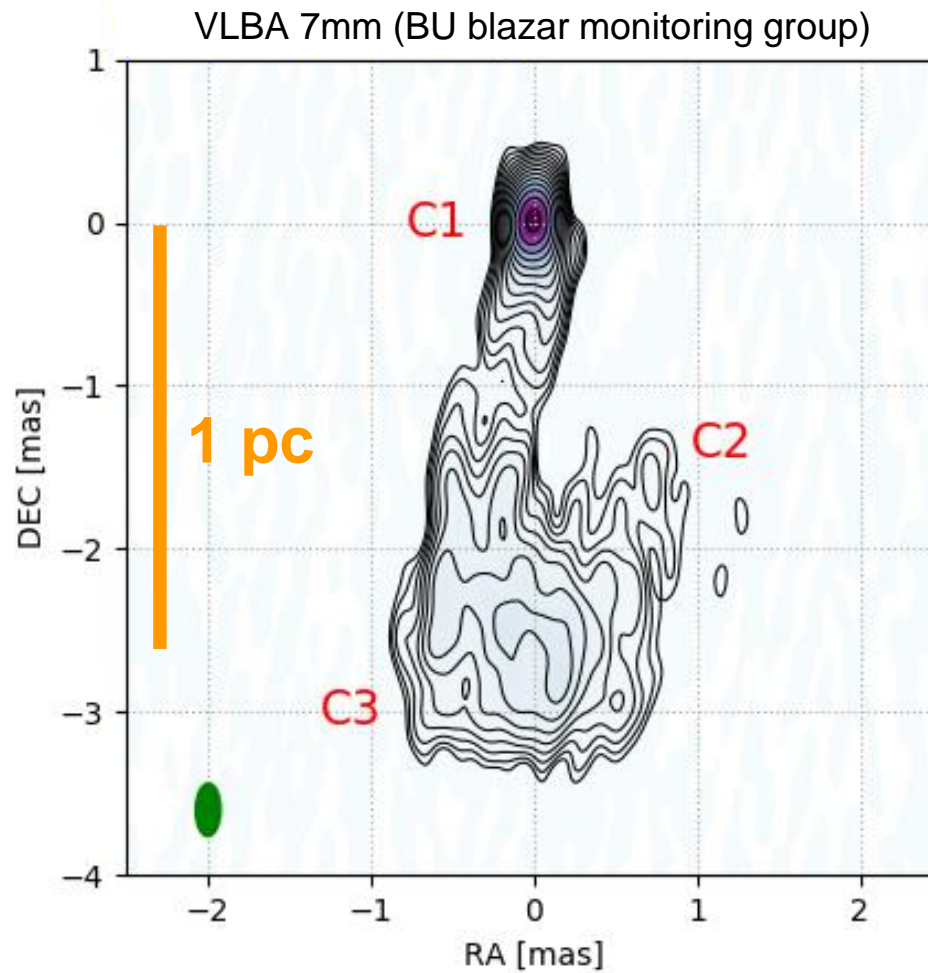
- This presentation has received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement No 730562 [RadioNet]



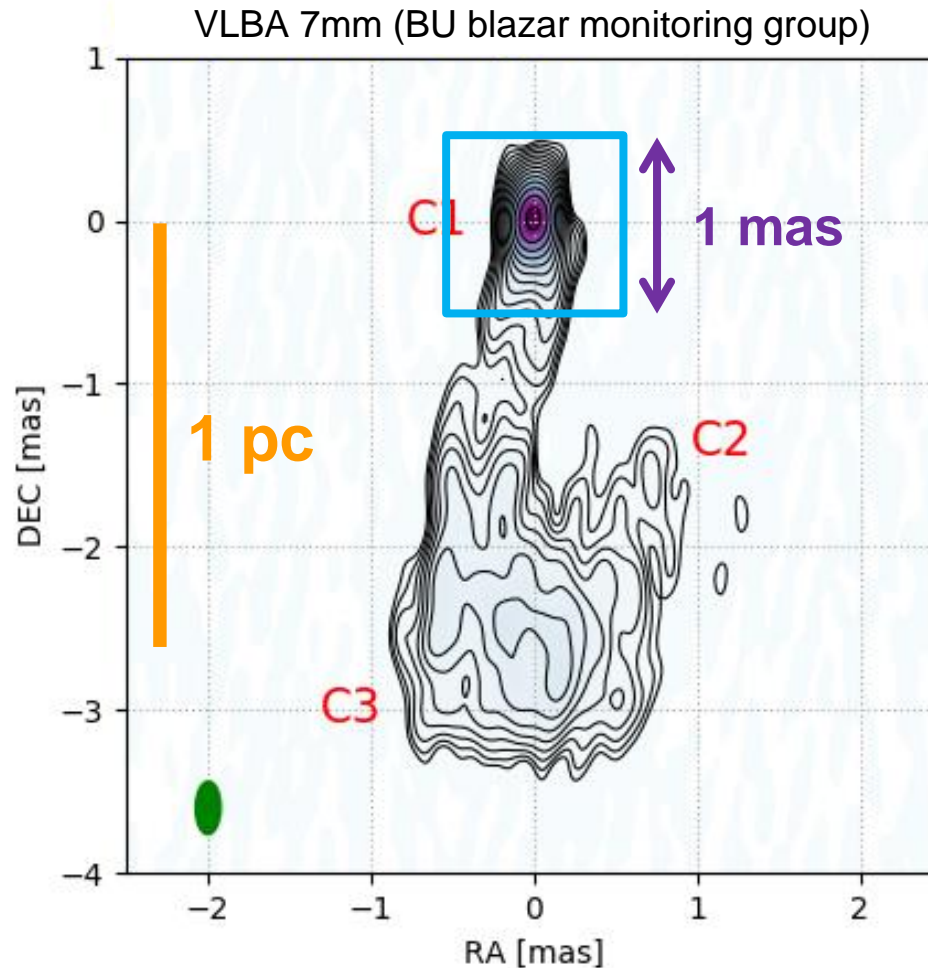
Introduction

- Probing jet launching mechanism (BZ vs BP) by direct imaging of jet launching region (jet morphology)
 - Highest possible angular resolution
 - Nearby target
 - High observation frequency
- **3C84** is one of the best target sources

Angular scale

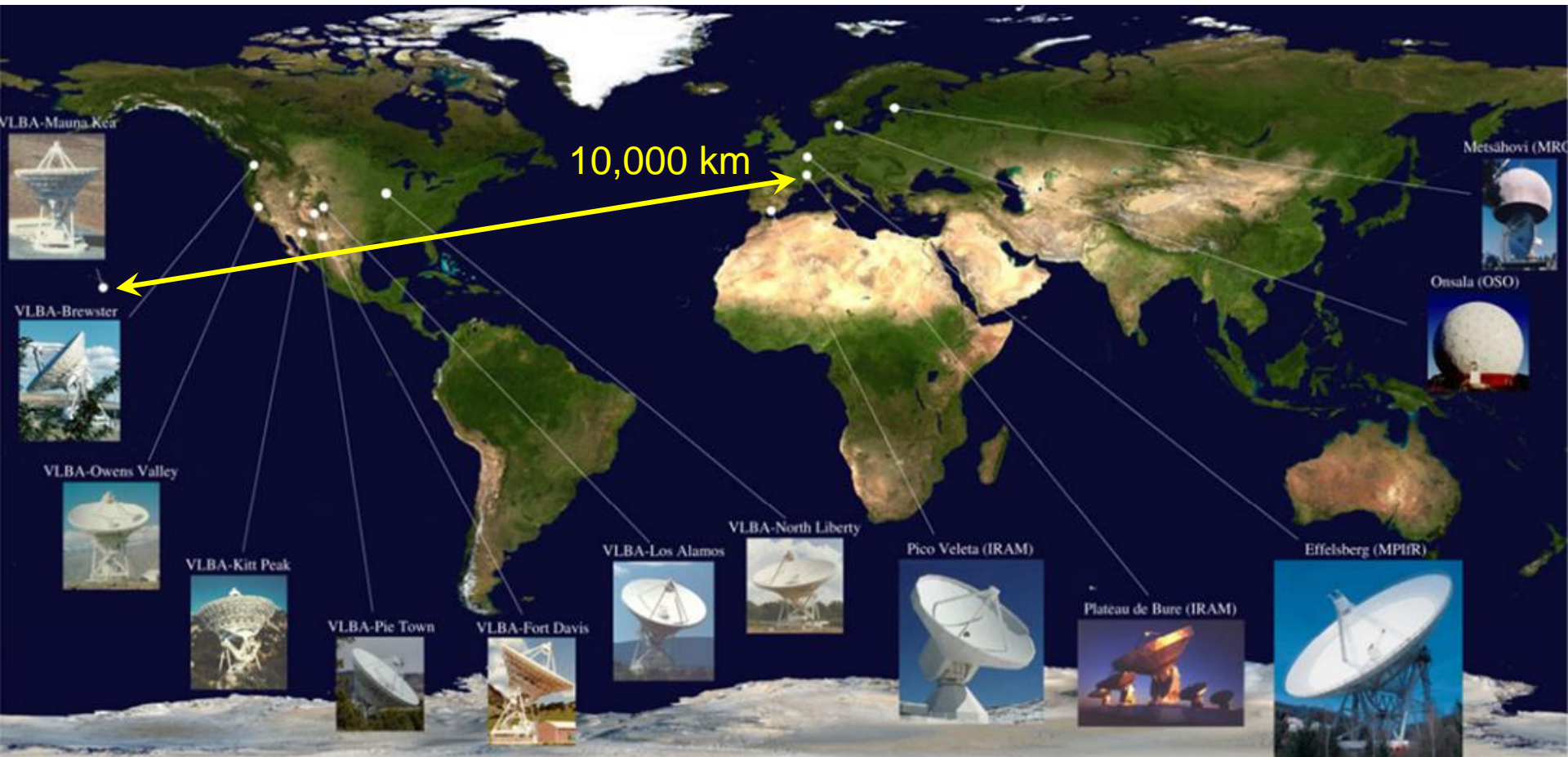


Angular scale



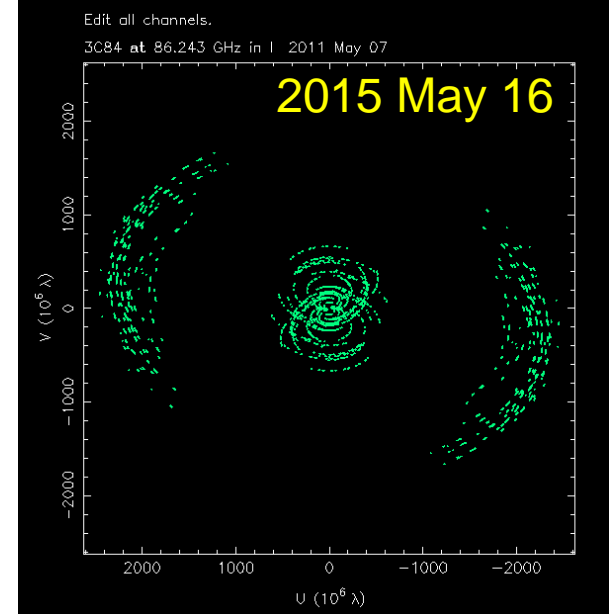
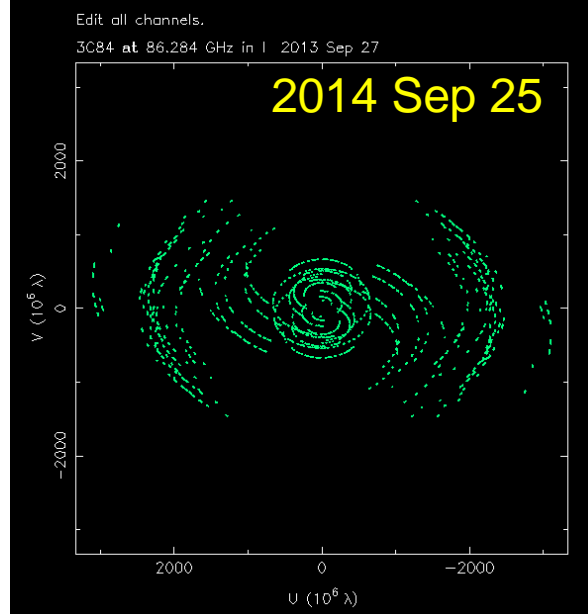
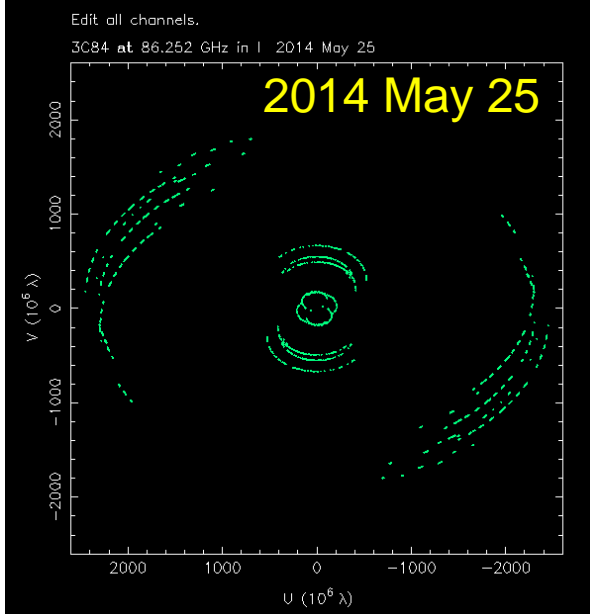
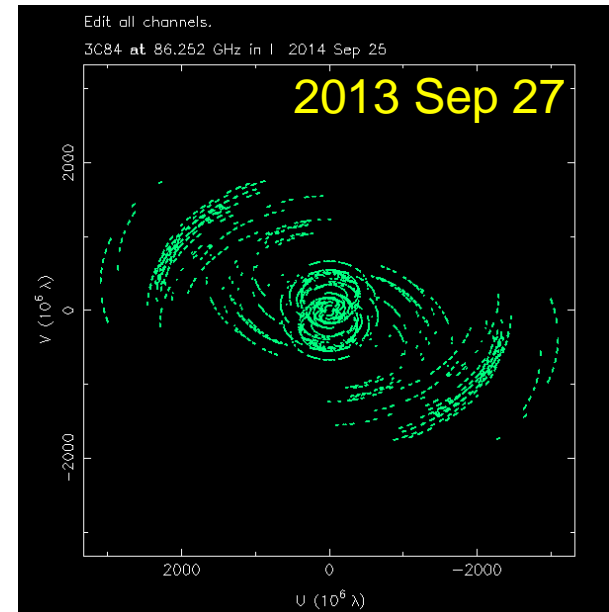
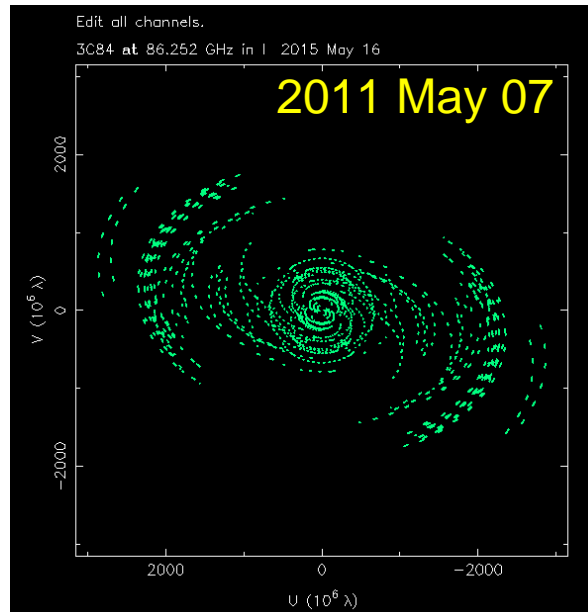
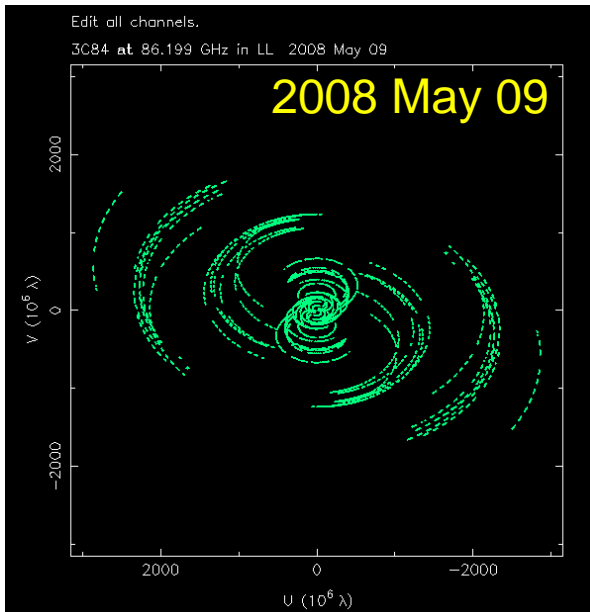
$1 \text{ mas} = 0.36 \text{ pc} \approx 11000 R_S$
 $M_{\bullet} = 3.2 \times 10^8 M_{\odot}$ (Park & Trippe 2017)
 $H_0 = 70 \text{ km/s/Mpc}$

Observations - GMVA



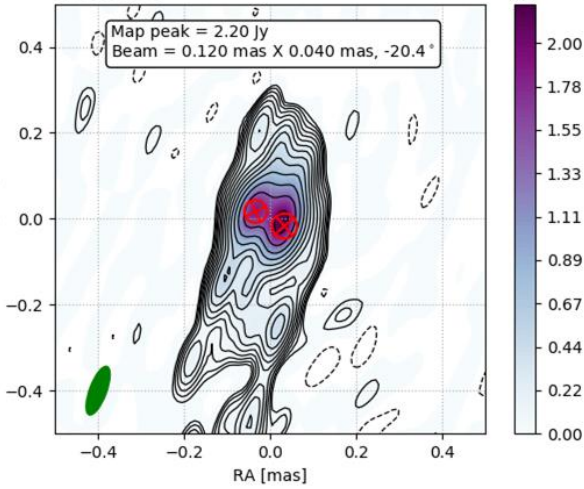
14+ telescopes (+GBT / +KVN)
maximum baseline length : ~10,000 km
Angular resolution : **50 ~ 70 μ as**
Operating at **86 GHz (3mm)**

Observations – uv coverage

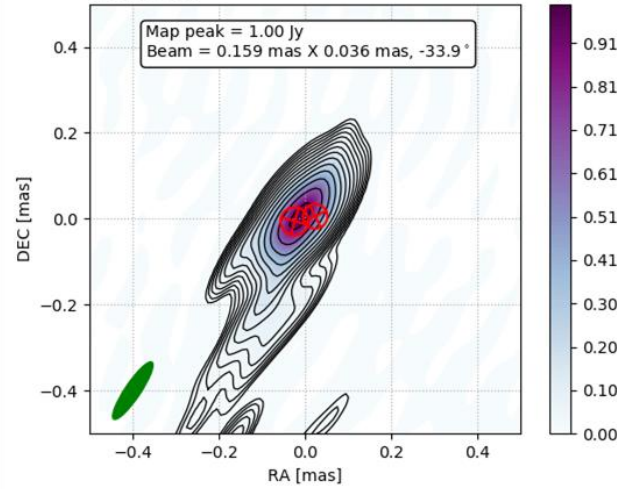


Consistent double nuclear structure in all 6 epochs

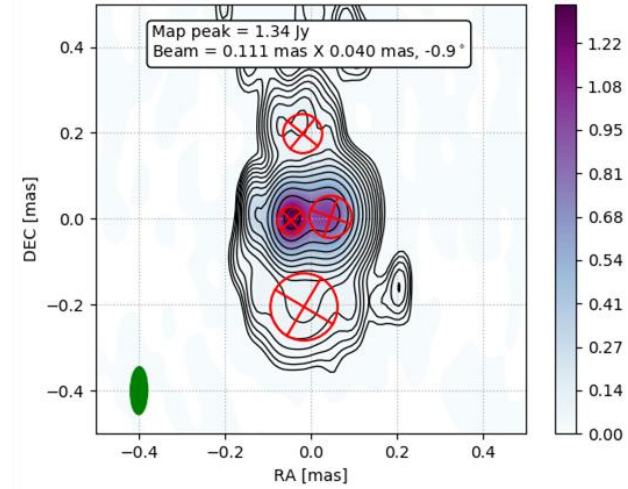
2008-05-09



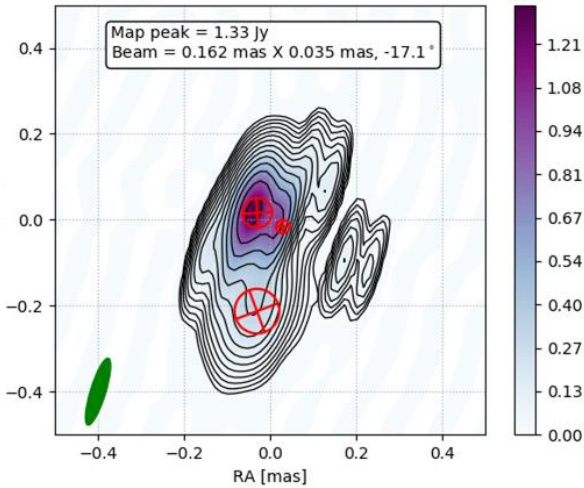
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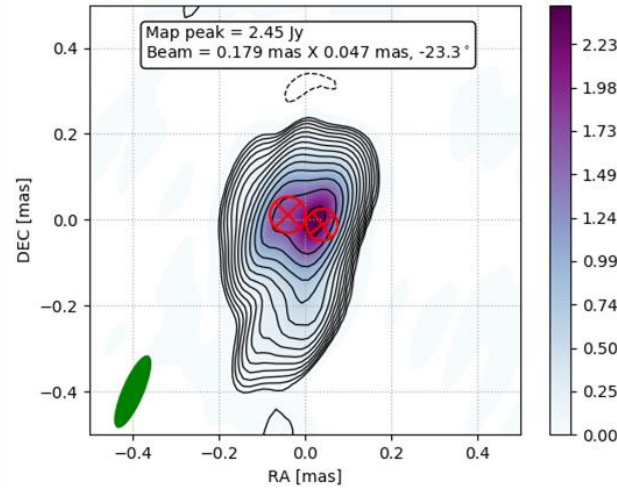
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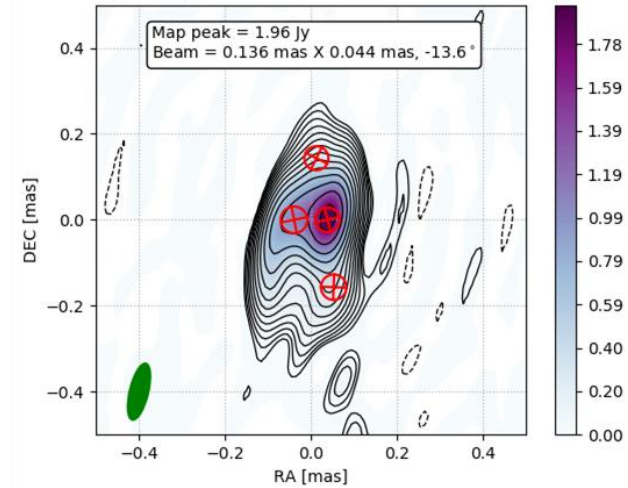
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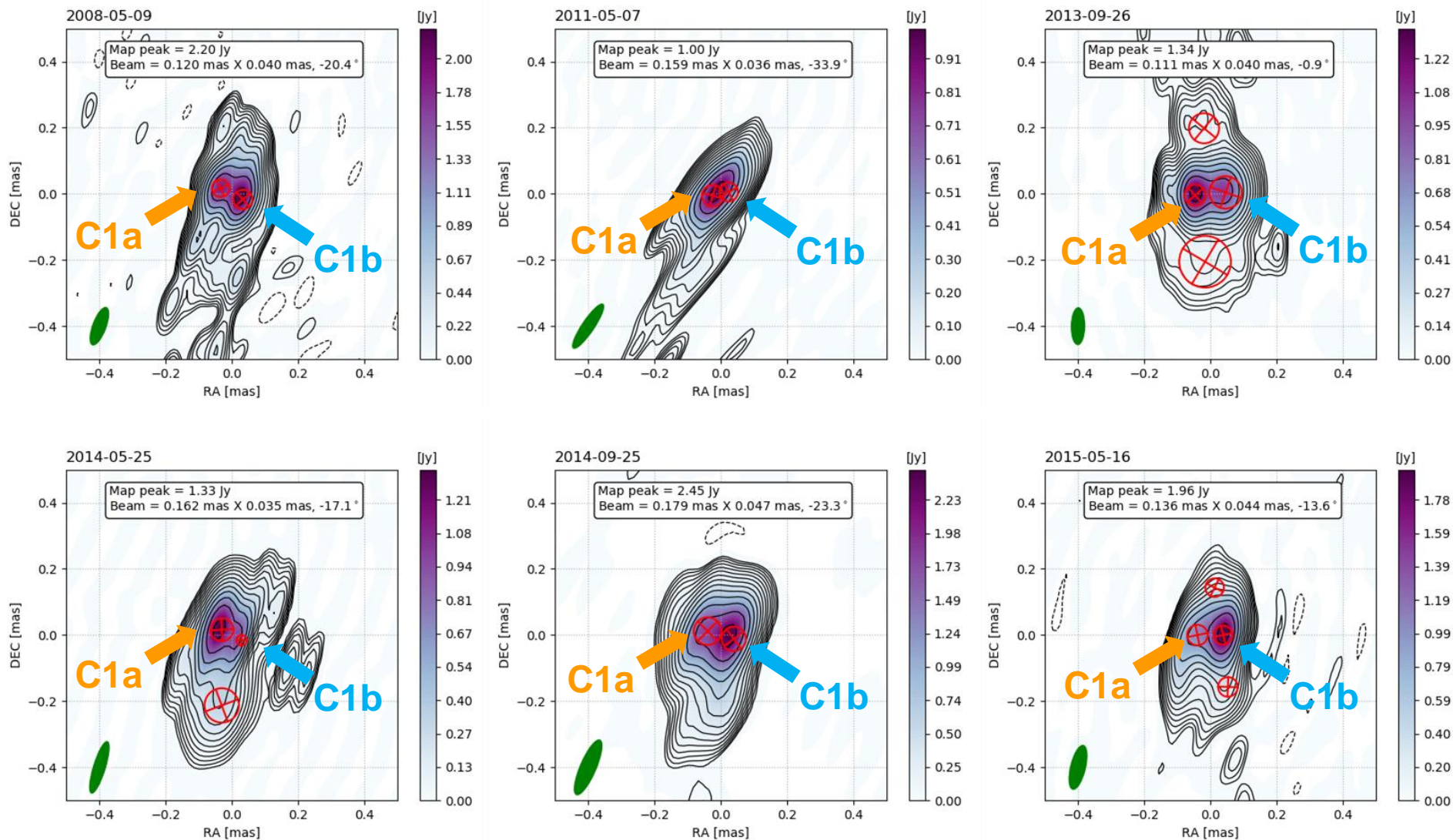
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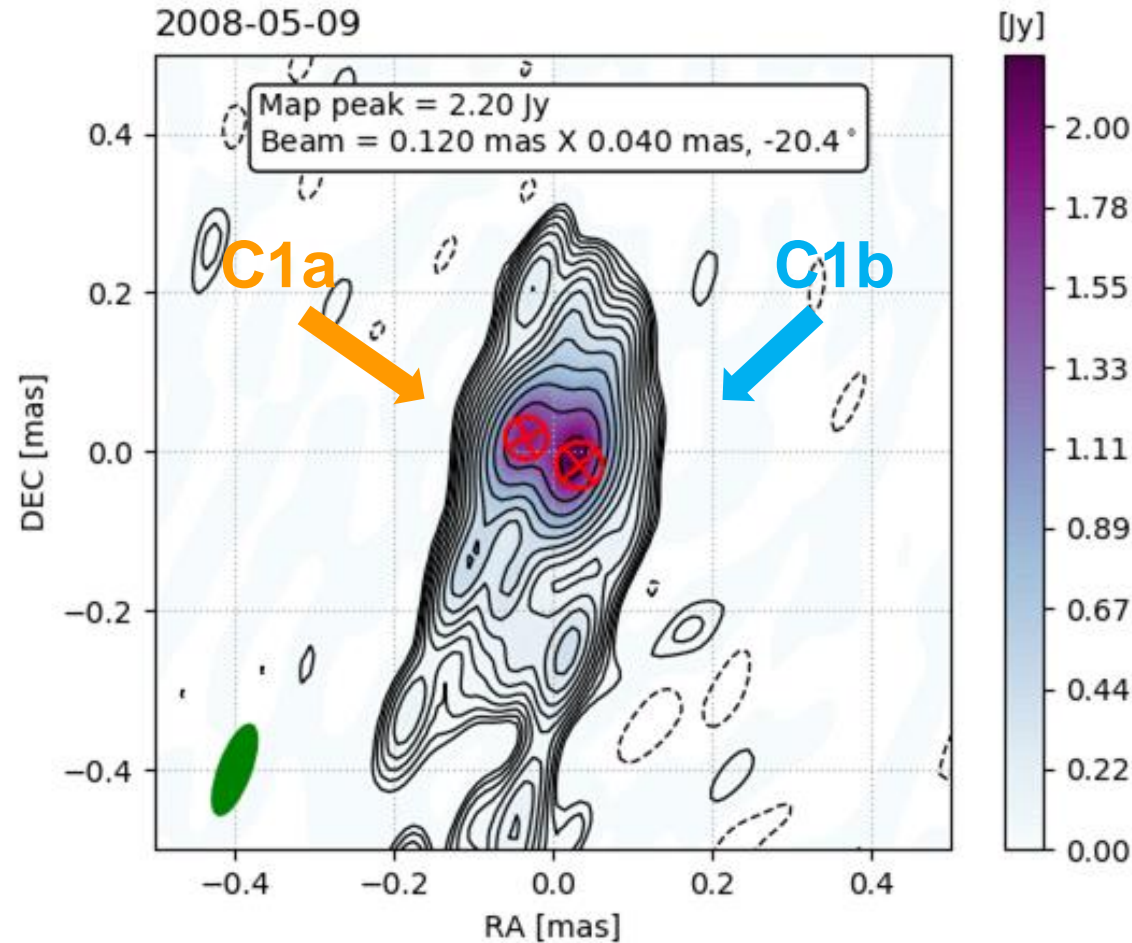
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Consistent double nuclear structure in all 6 epochs

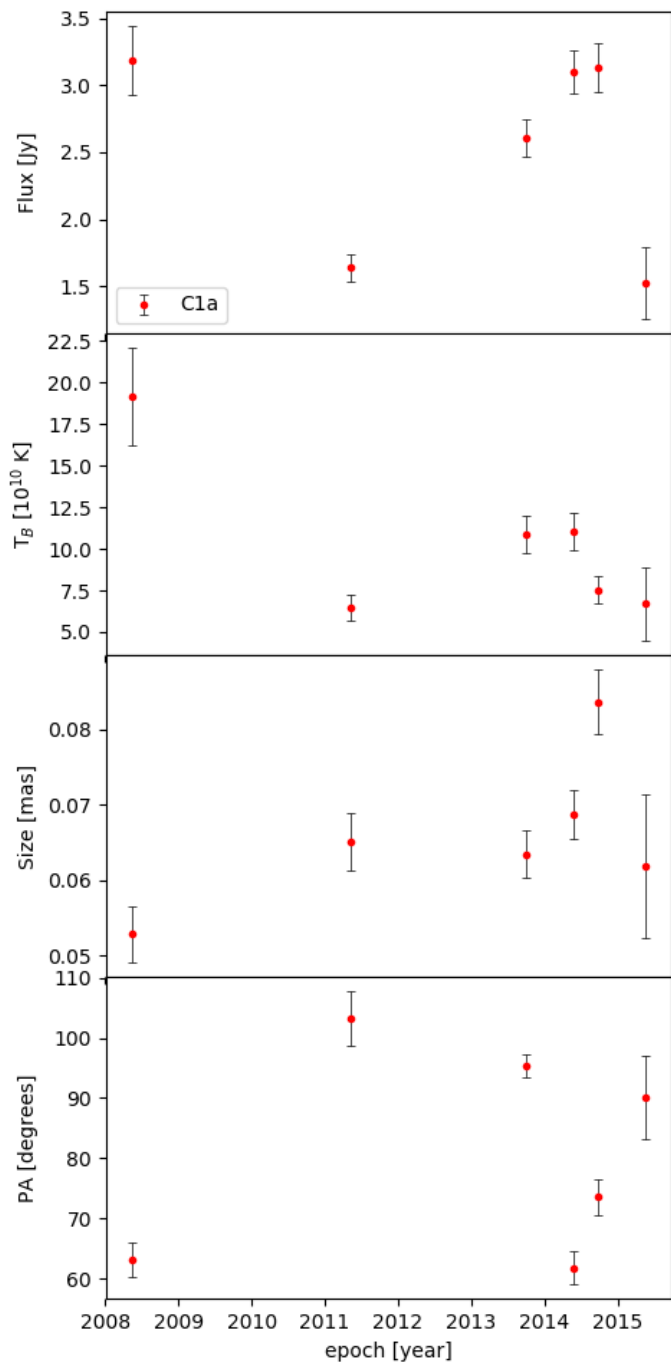


Consistent double nuclear structure in all 6 epochs

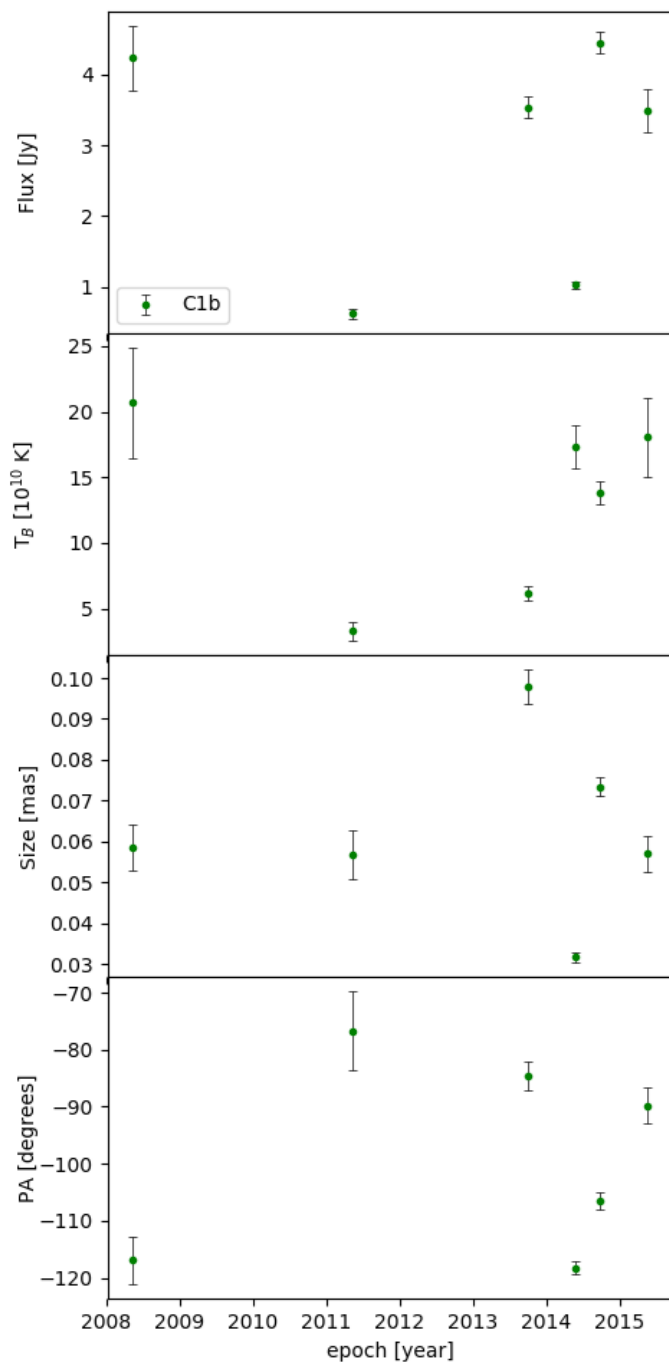


- No significant motion over 8 years
- Separation $\sim 70 \mu\text{as}$
- Brightness temperature
C1a : 1.0×10^{11} K
C1b : 1.3×10^{11} K
- Continued to Limb-brightened jet structure

● C1a



● C1b



- Distance between C1a and C1b
~800 R_S (~1 light-month, for $M_{\text{BH}} = 3.2 \times 10^8 M_{\odot}$)
- If C1a + C1b is jet base, we have Blandford–Payne mechanism at work (Blandford–Znajek requires $<10 R_S$)

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- But the size is probably even too large for an accretion disk
- Accretion disk size vs. black hole mass (Morgan+ 2010)

$$\log\left(\frac{R_{2500}}{cm}\right) = (15.78 \pm 0.12) + (0.80 \pm 0.17)\log\left(\frac{M_{BH}}{10^9 M_\odot}\right)$$

- Expected for 3C84 : ~54 R_S

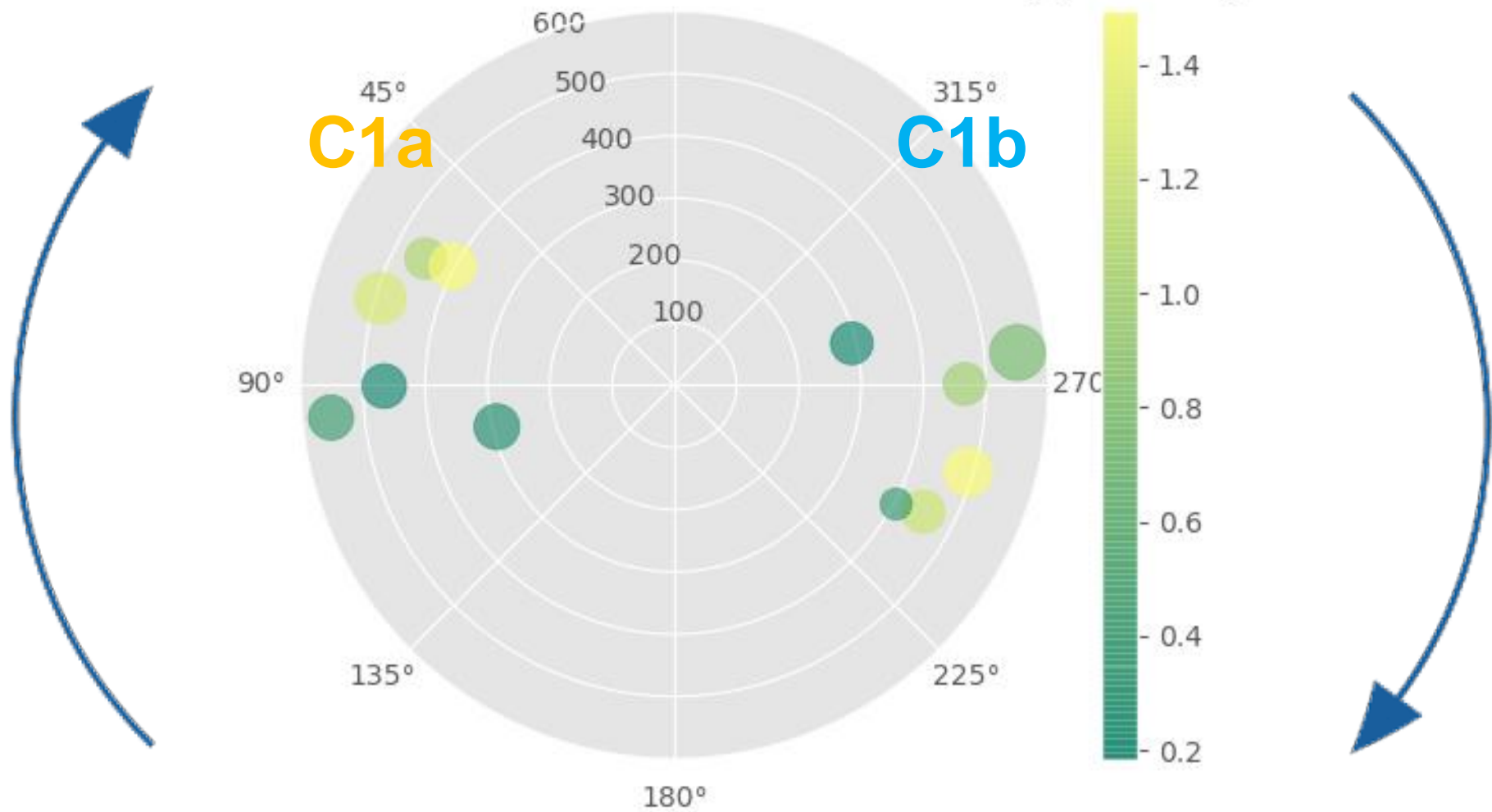
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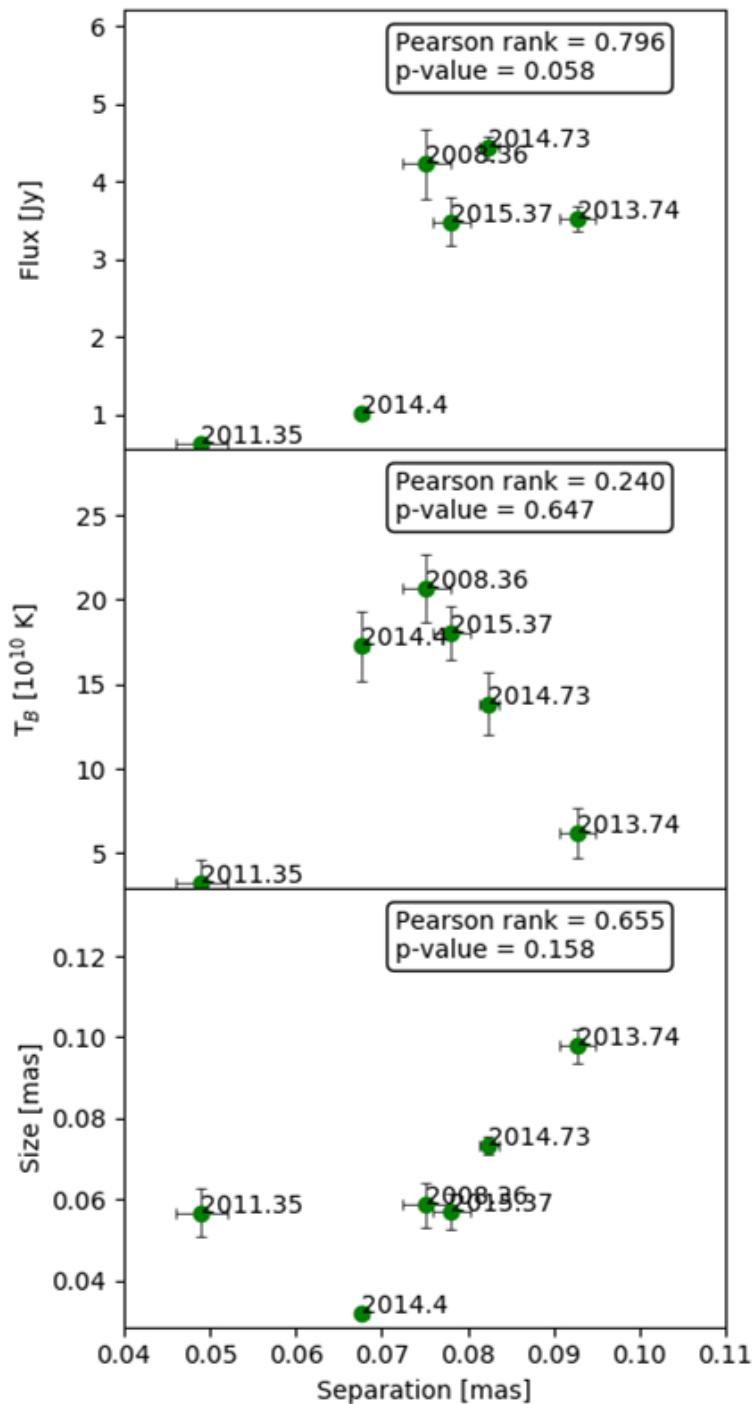
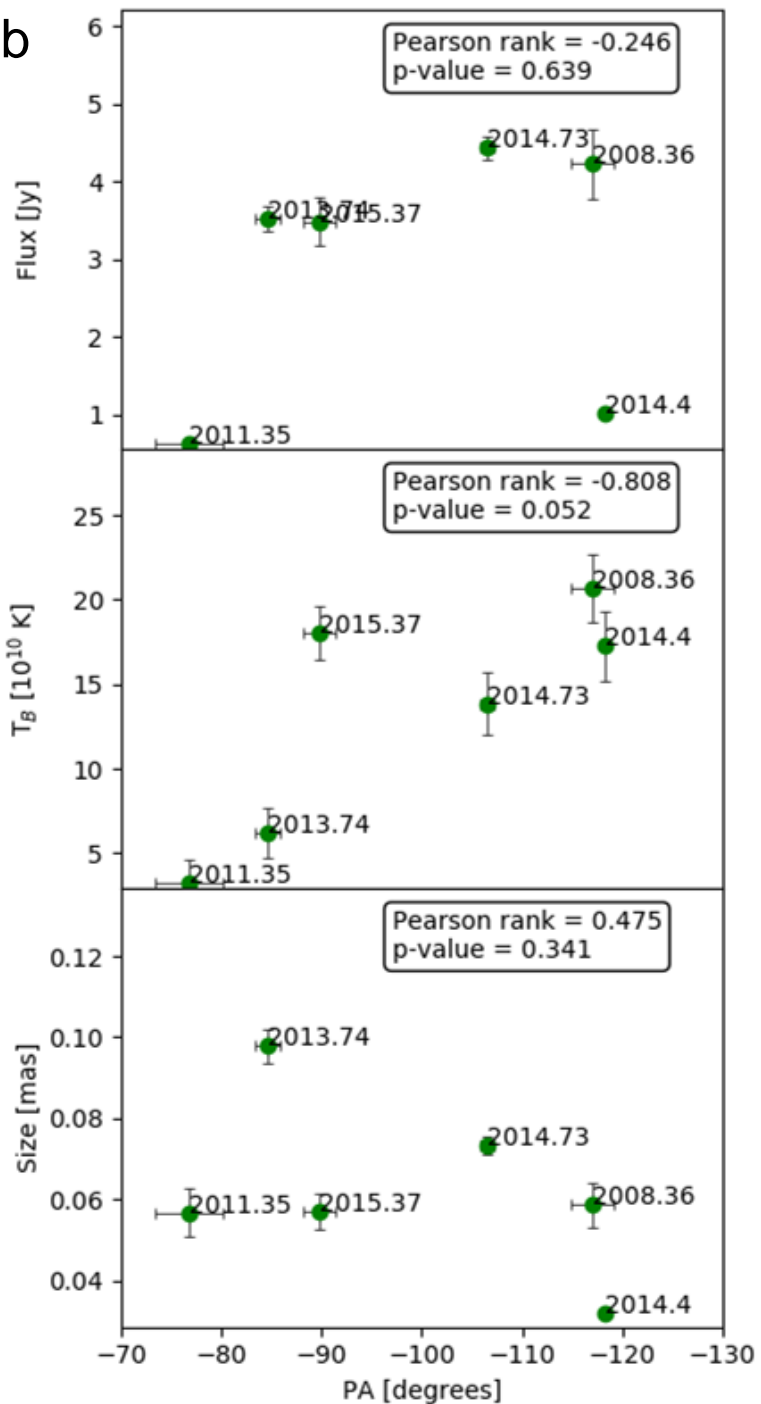
- Expected for 3C84 : ~54 R_S
- High brightness temperature ($>10^{11}$ K) indicates non-thermal emission

Relative locations: $r: R_S$

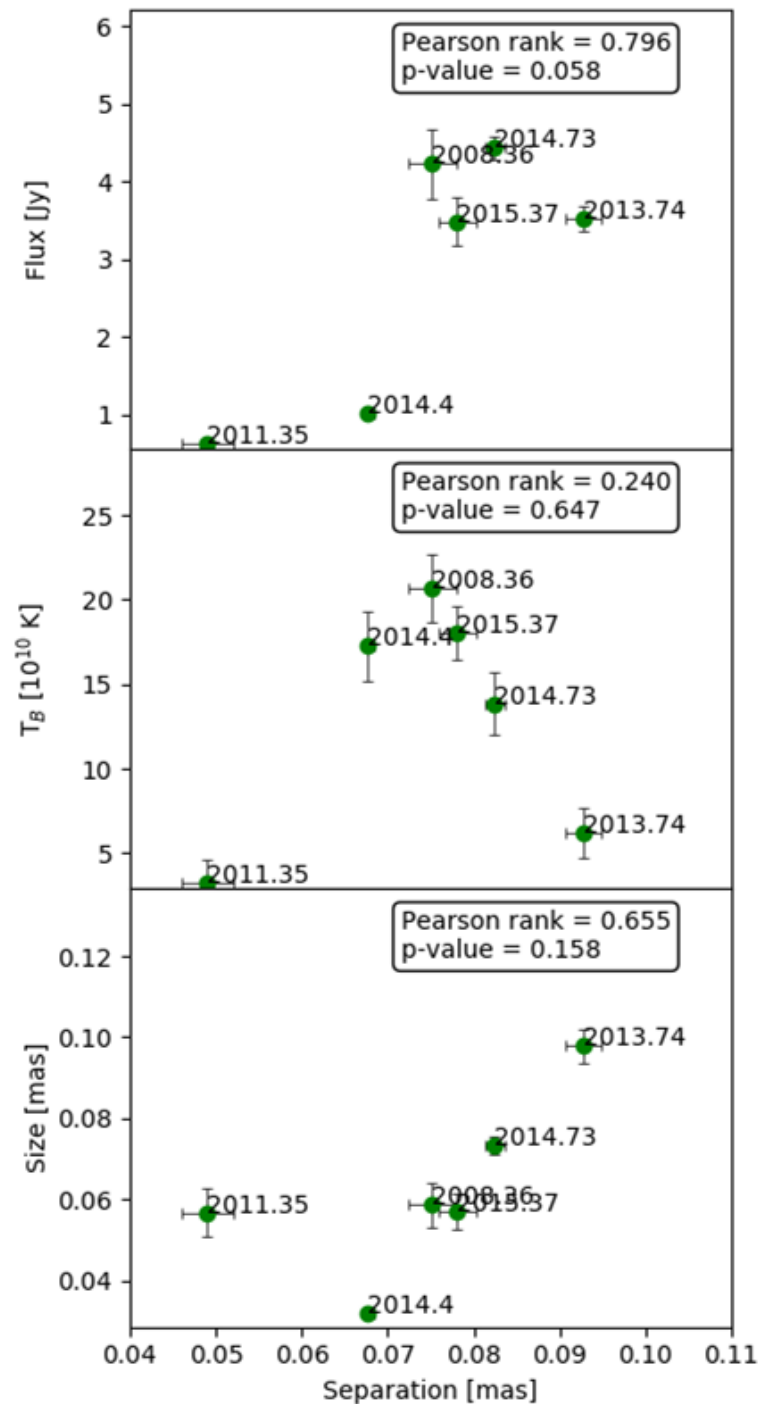
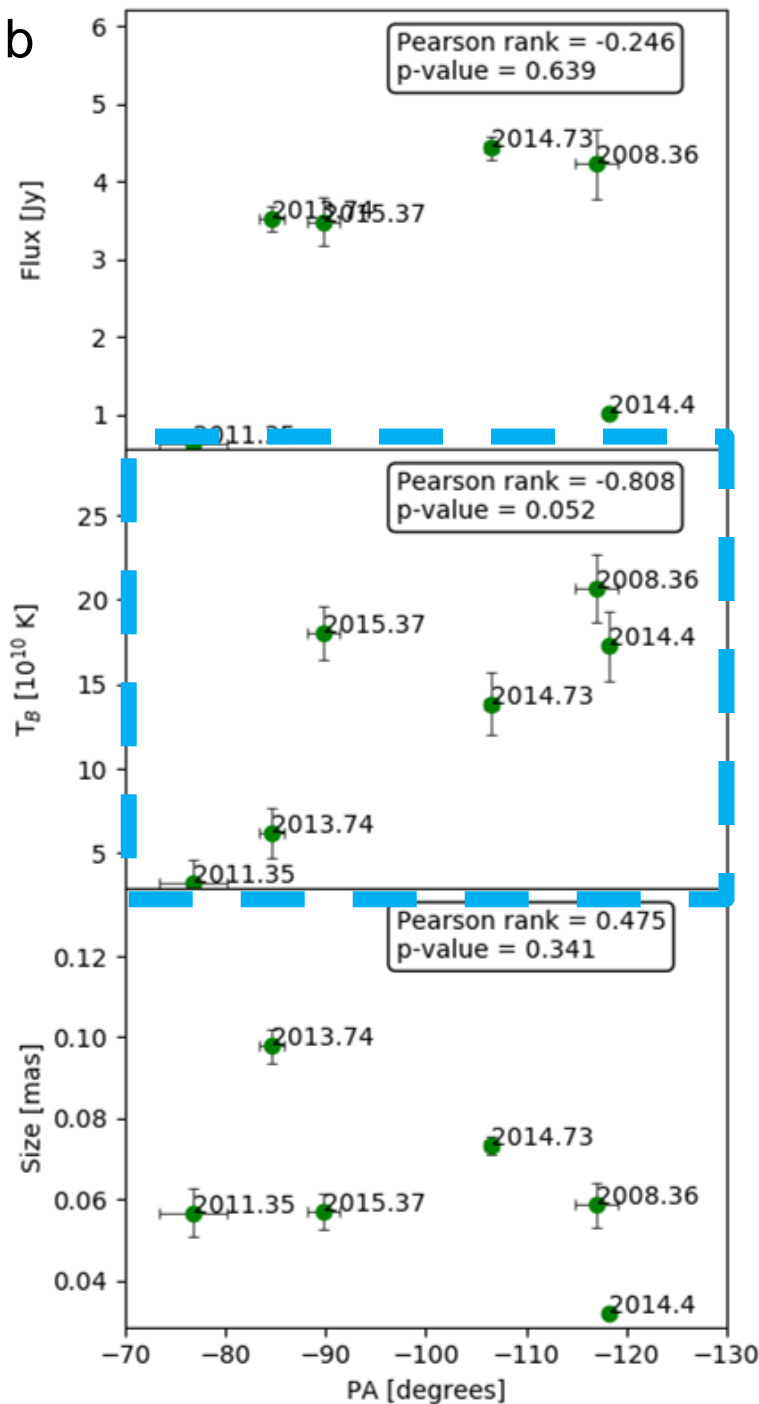
$T_B [\times 10^{10} \text{ K}]$



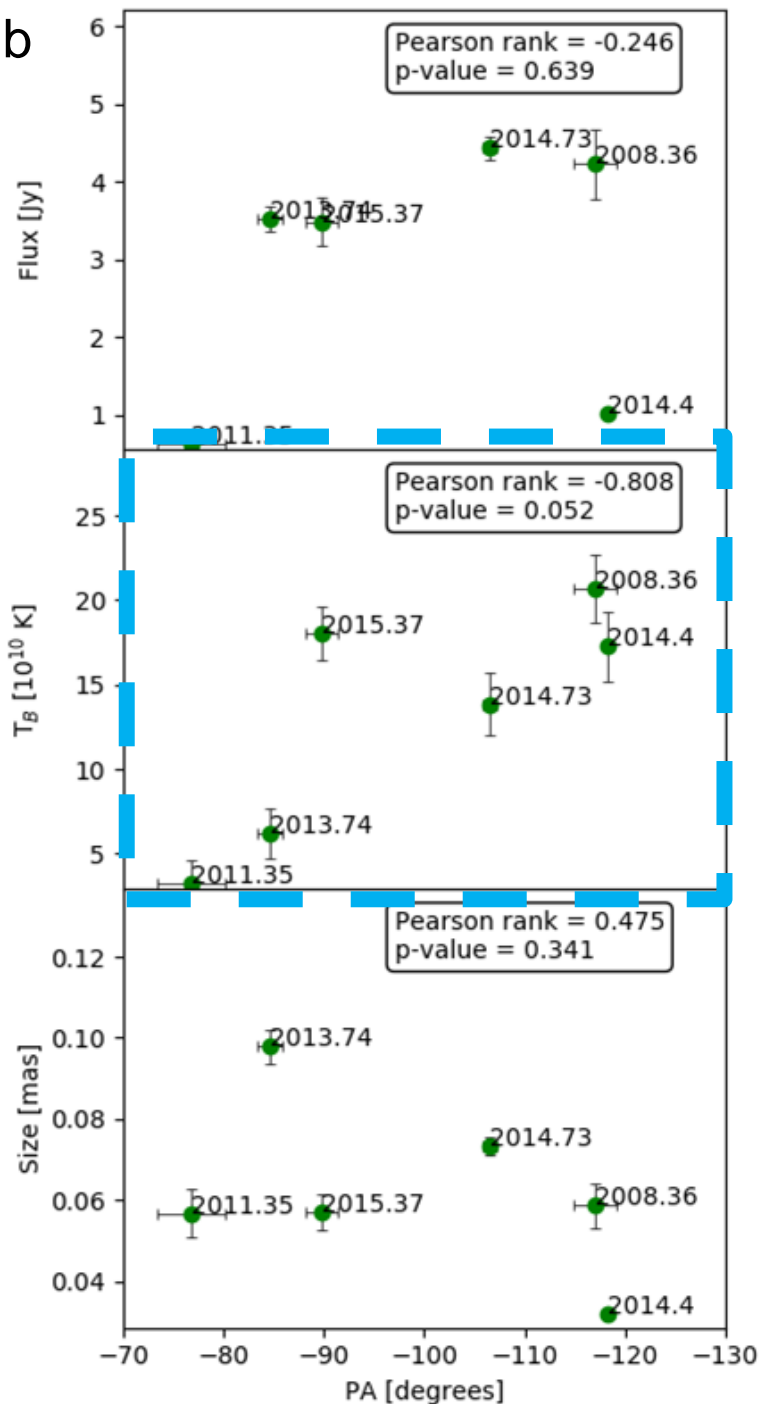
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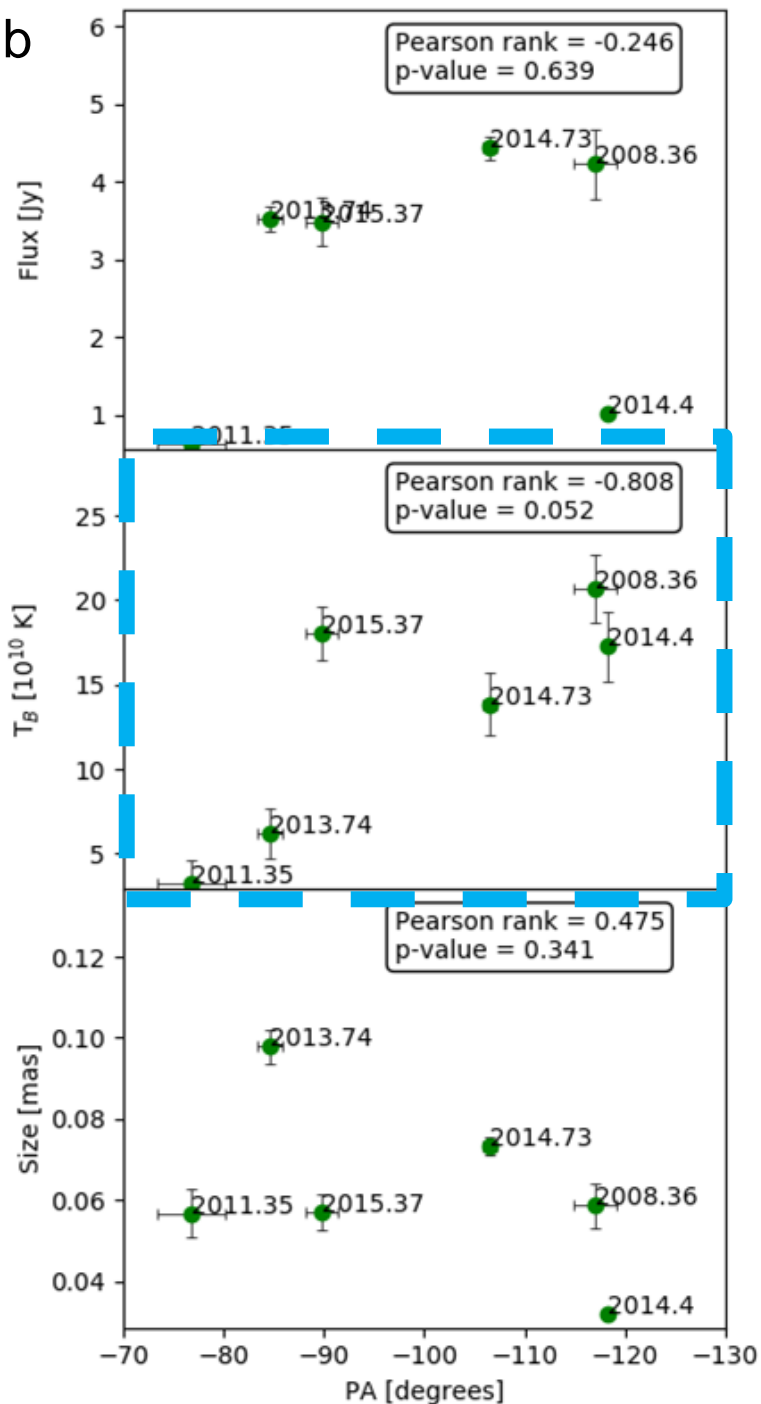


- T_B – PA and S_V – PA correlation
- T_B varies by factor of ~ 6

→ Emitters moving on a helical path

- Possible physical processes
 1. Doppler boosting
 2. Intrinsic evolution of the jet plasma

● C1b



- T_B – PA and S_V – PA correlation
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→ Emitters moving on a helical path

- Possible physical processes
 1. Doppler boosting
 2. Intrinsic evolution of the jet plasma

$$T_B^{obs} \leq 2 \times 10^{11} K$$

Equipartition limit (Singal 2009)

$$T_B^{em} \leq \sim 10^{11} K$$

$$\delta \approx 1 \quad \longrightarrow \quad \beta \ll 1$$

Intrinsic evolution of jet plasma

- Assuming all emission is synchrotron radiation
- No correlation of flux with time
- Multiple individual emitters cooling down rapidly

Intrinsic evolution of jet plasma

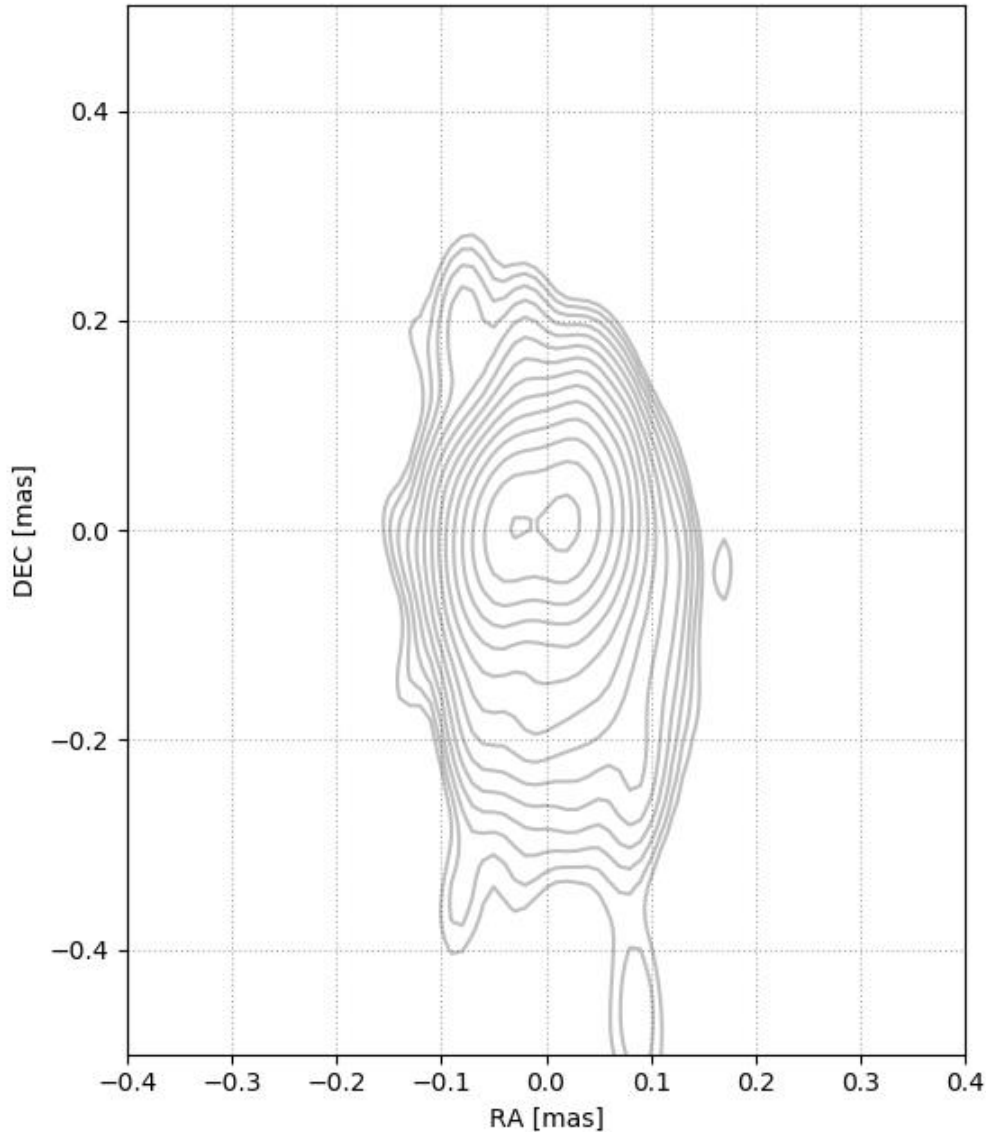
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- The cooling time scale :

$$\tau_{cool} = 7.74 \left[\frac{\delta}{1+z} \right]^{-1} B^{-2} \gamma^{-1} \text{ seconds}$$

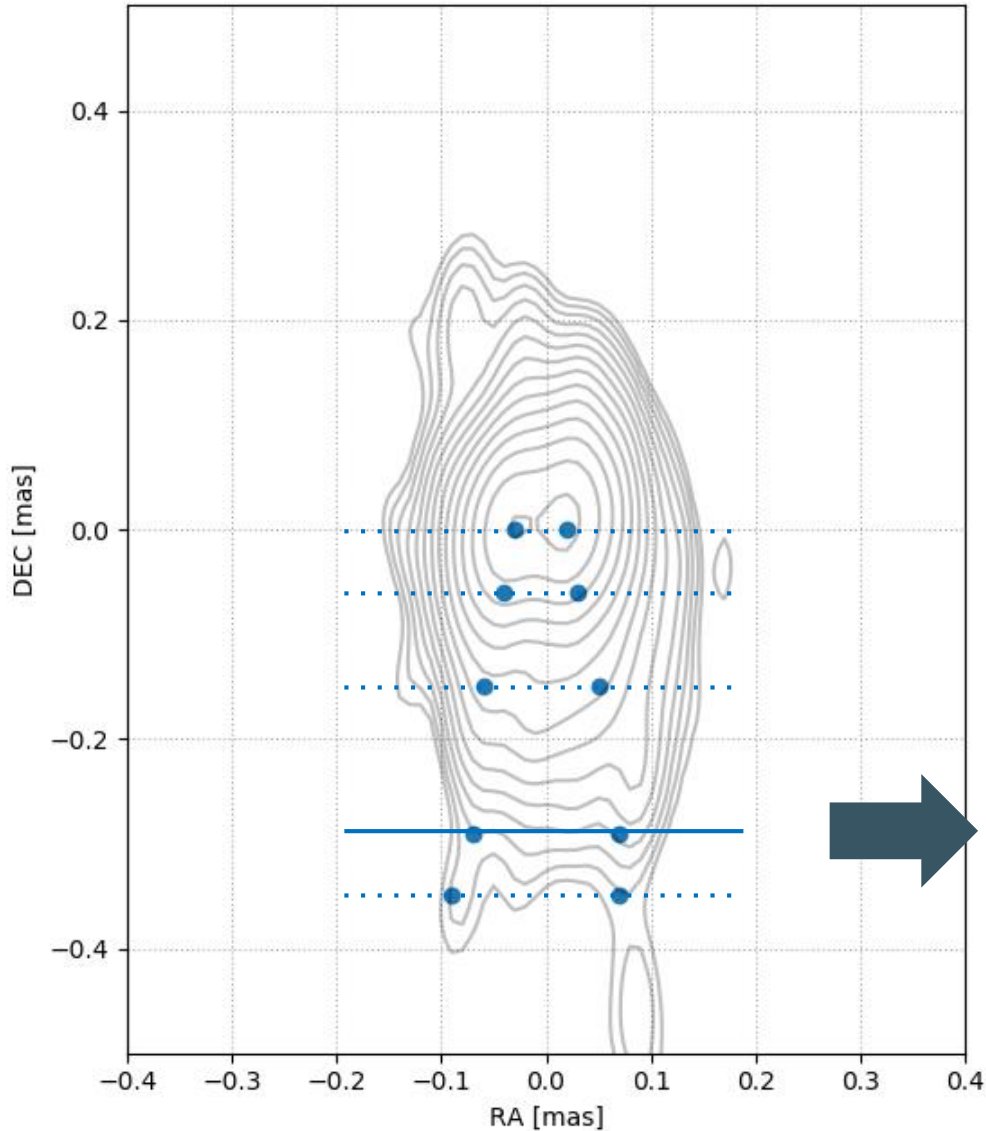
- $\delta \approx 1, B \approx 10 \mu T, \gamma \approx 10000, z = 0.0176$
- ~3 months
- Typical blazar-like value (Hodgson+ 2016)

Where is the Black Hole?

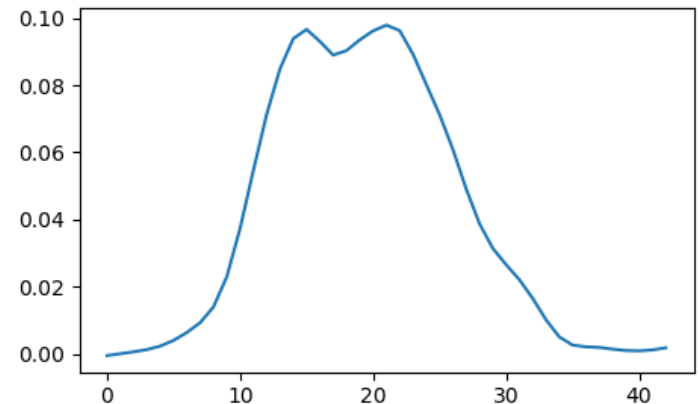


- Jet profile using multiple slices on stacked map
- Found local maxima

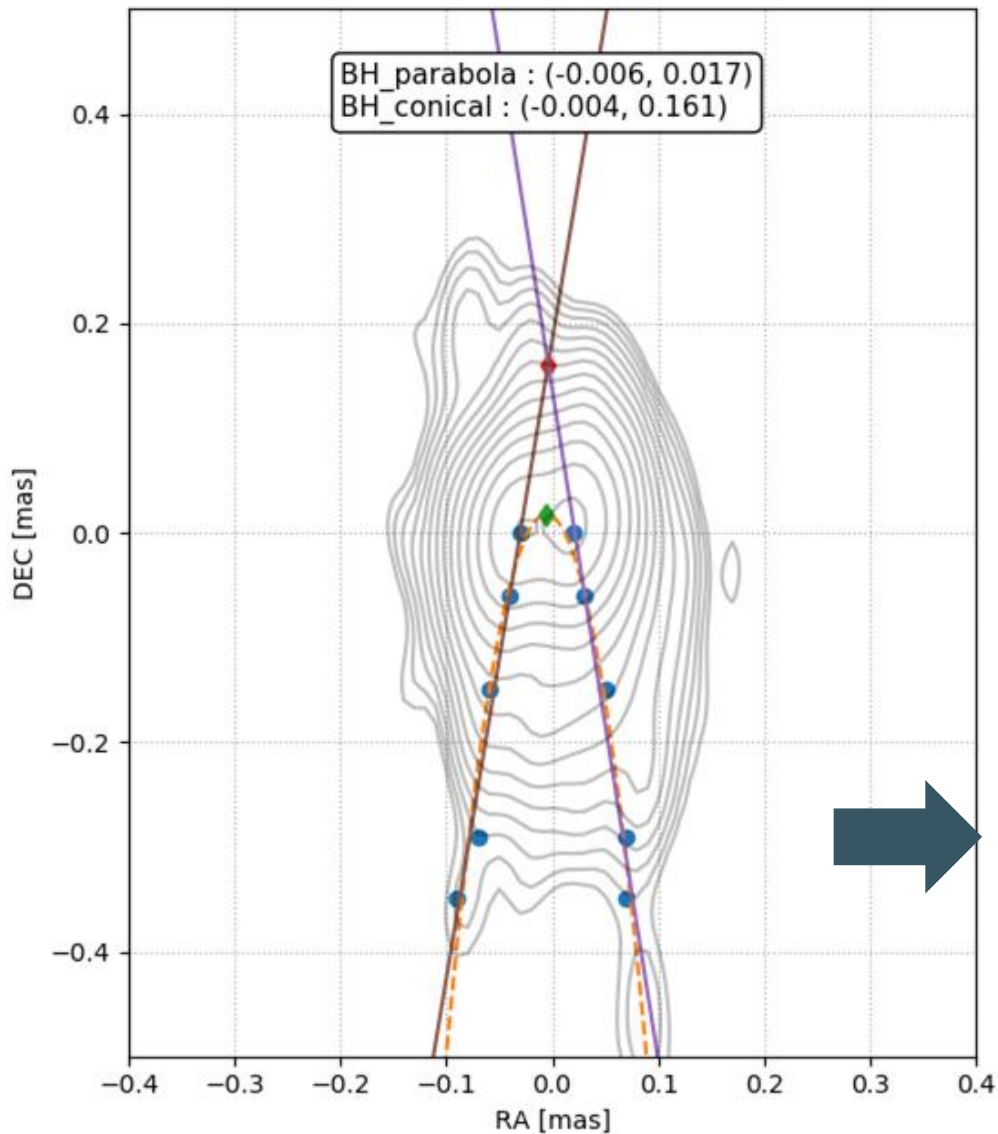
Where is the Black Hole?



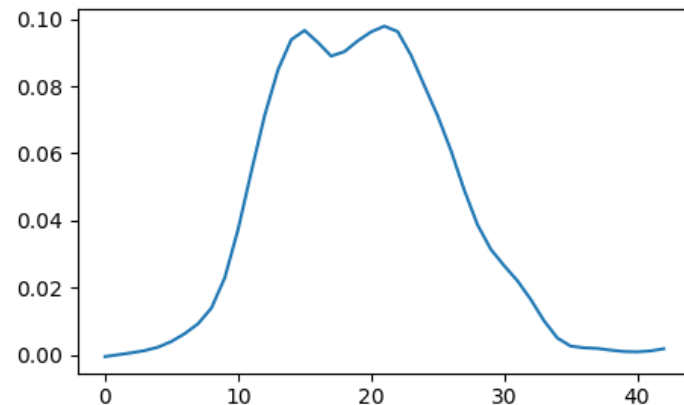
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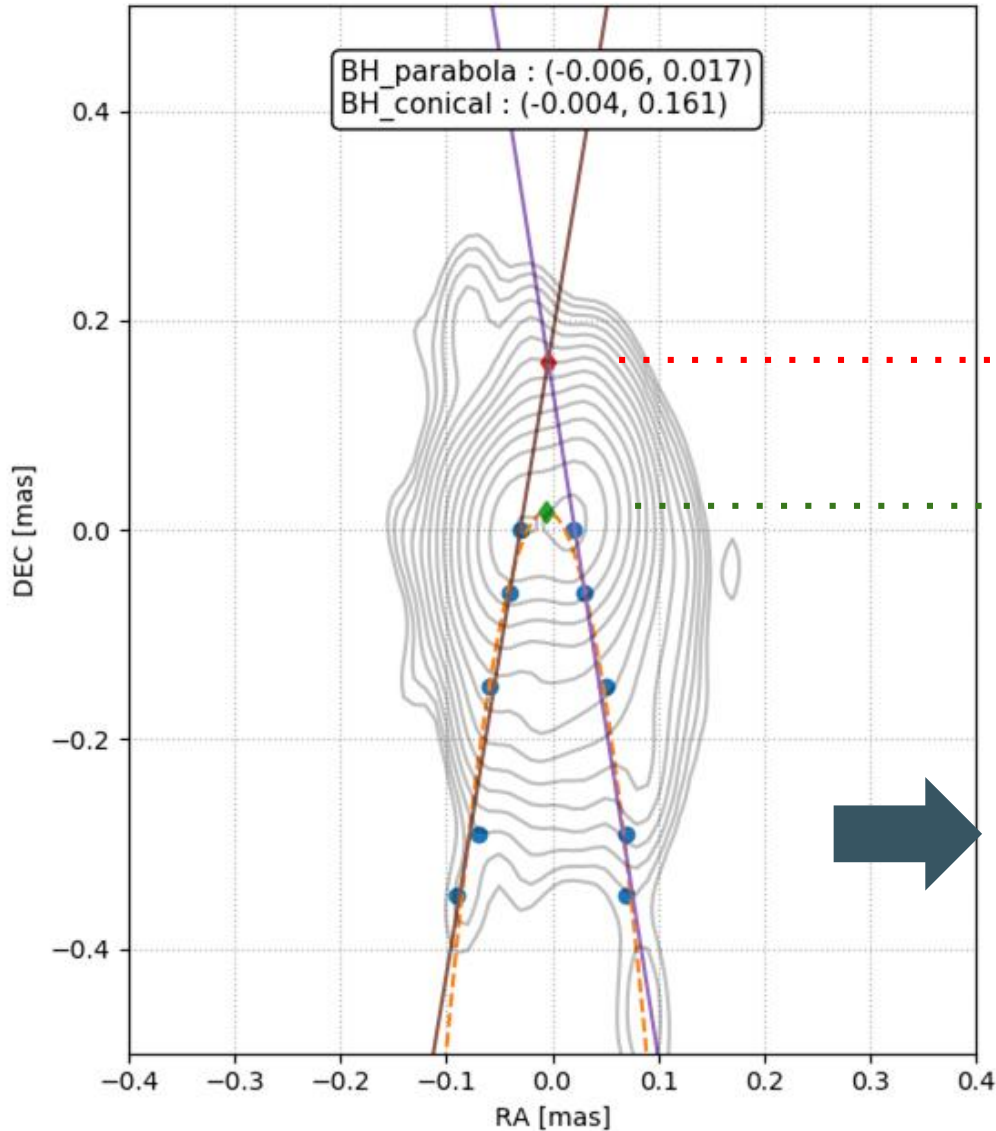
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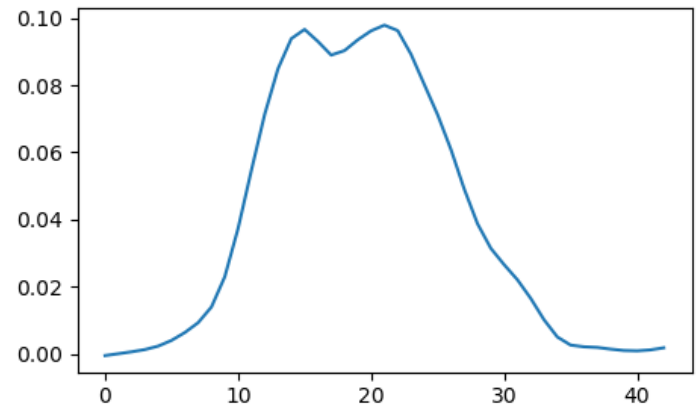
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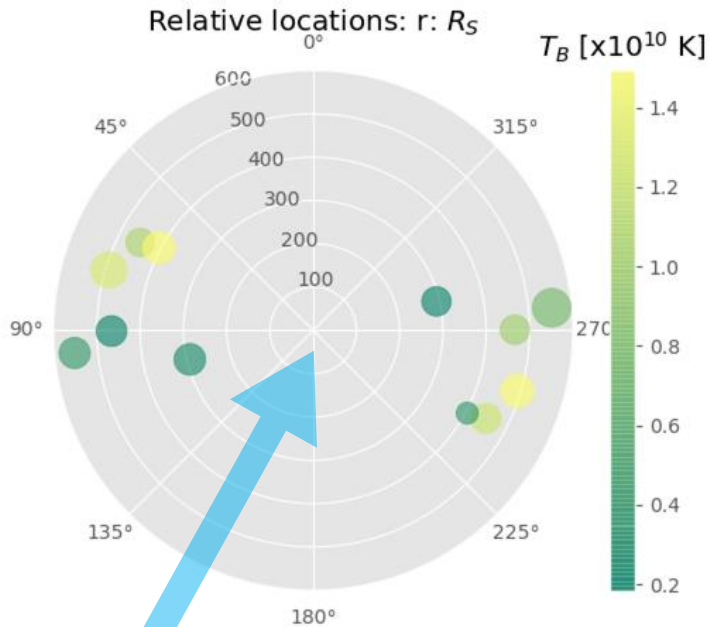
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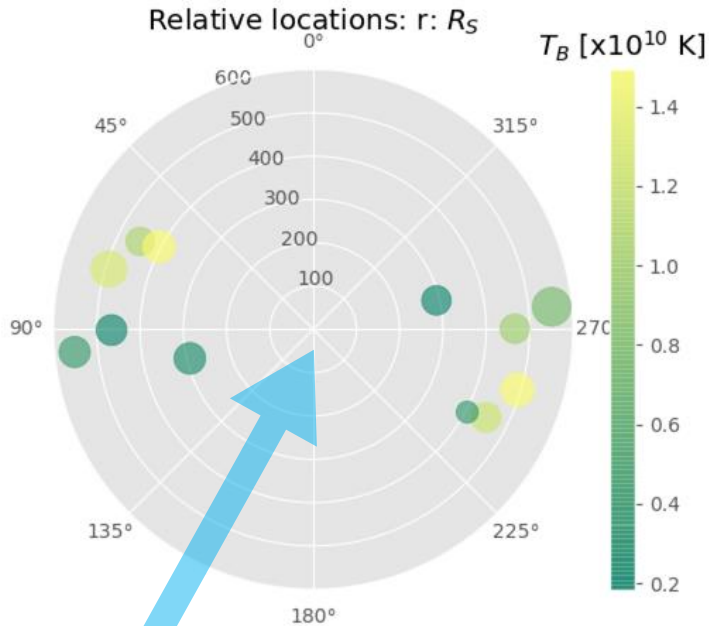
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- If jet plasma is expanding and cooling, SMBH must be between C1a and C1b
- Viewing angle in C1 $< 45^\circ$?
- Parabolic fit is more consistent

?

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Abdo et al. (2009)
 $\sim 25^\circ$
 γ -ray SED fitting

Fujita & Nagai (2017)
 $\sim 65^\circ$
Jet / counter-jet

?

- γ -ray emission from C1 region (Hodgson et al. 2018)
- Viewing angle changed from the nuclear region to the extended structure?

Summary

- An east-west oriented “double” nuclear structure in C1 region
- The brightness temperature of C1a and C1b, in the order of 10^{11}K and shows a trend of increasing brightness temperature to the north for C1a and to the south for C1b. This behavior is consistent with a helical expanding jet sheath.
- The behavior of the nuclear emission appears to be broadly consistent with that of a blazar.
- We placed limits of the true location of the SMBH assuming either a parabolic or conical jet to between $190 R_S$ and $1800 R_S$

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Thank you for listening!

